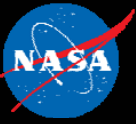




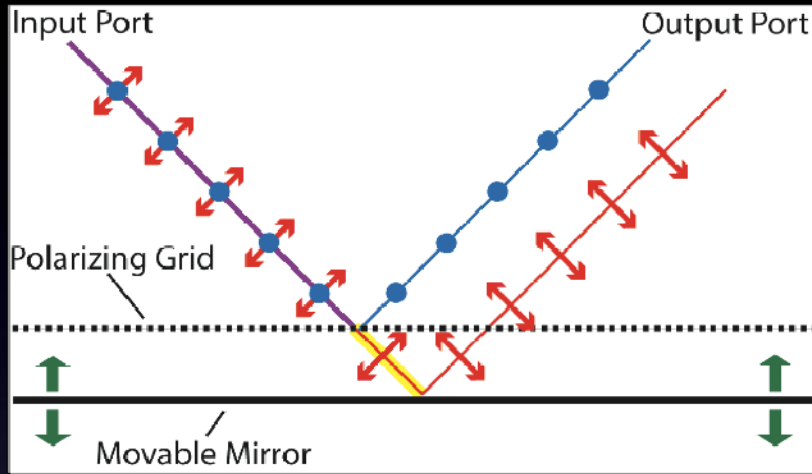
Goddard Space Flight Center

Variable-delay Polarization Modulators (VPMs)

David T. Chuss
NASA Goddard Space Flight Center
CMBPol Technology Workshop
Boulder, CO
August 27, 2008

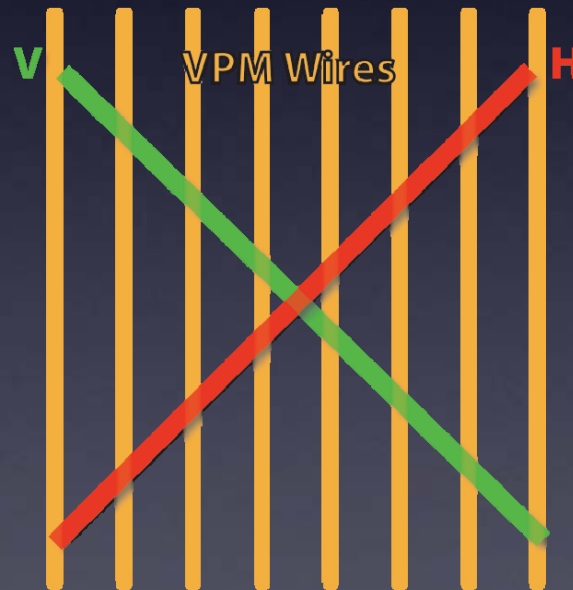
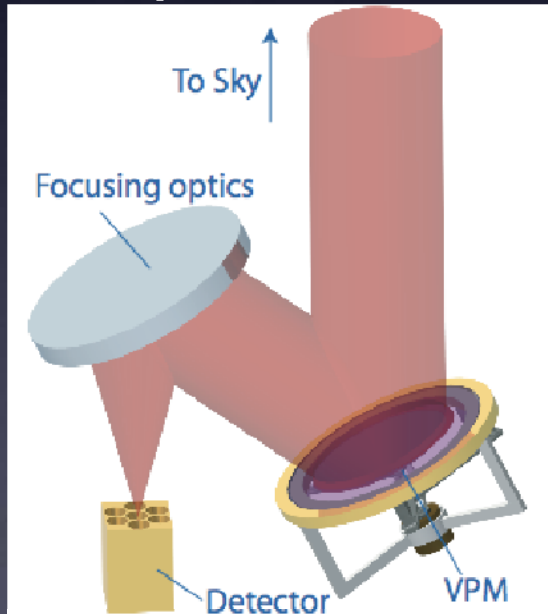


(VPMs)

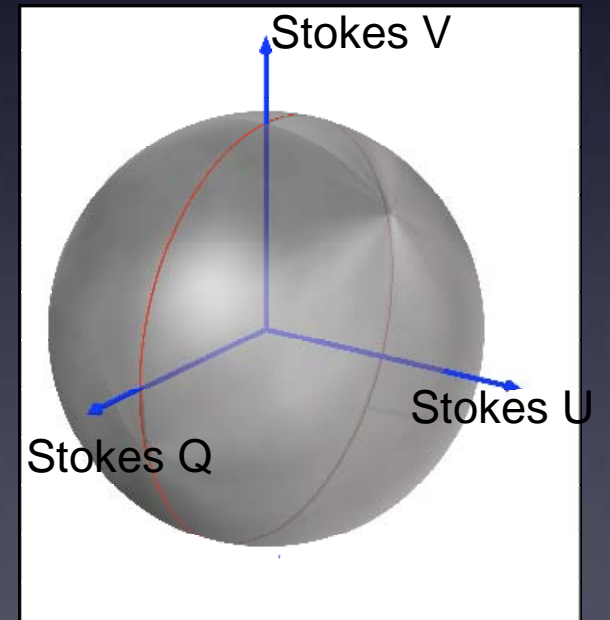


Operational Principle

Implementation

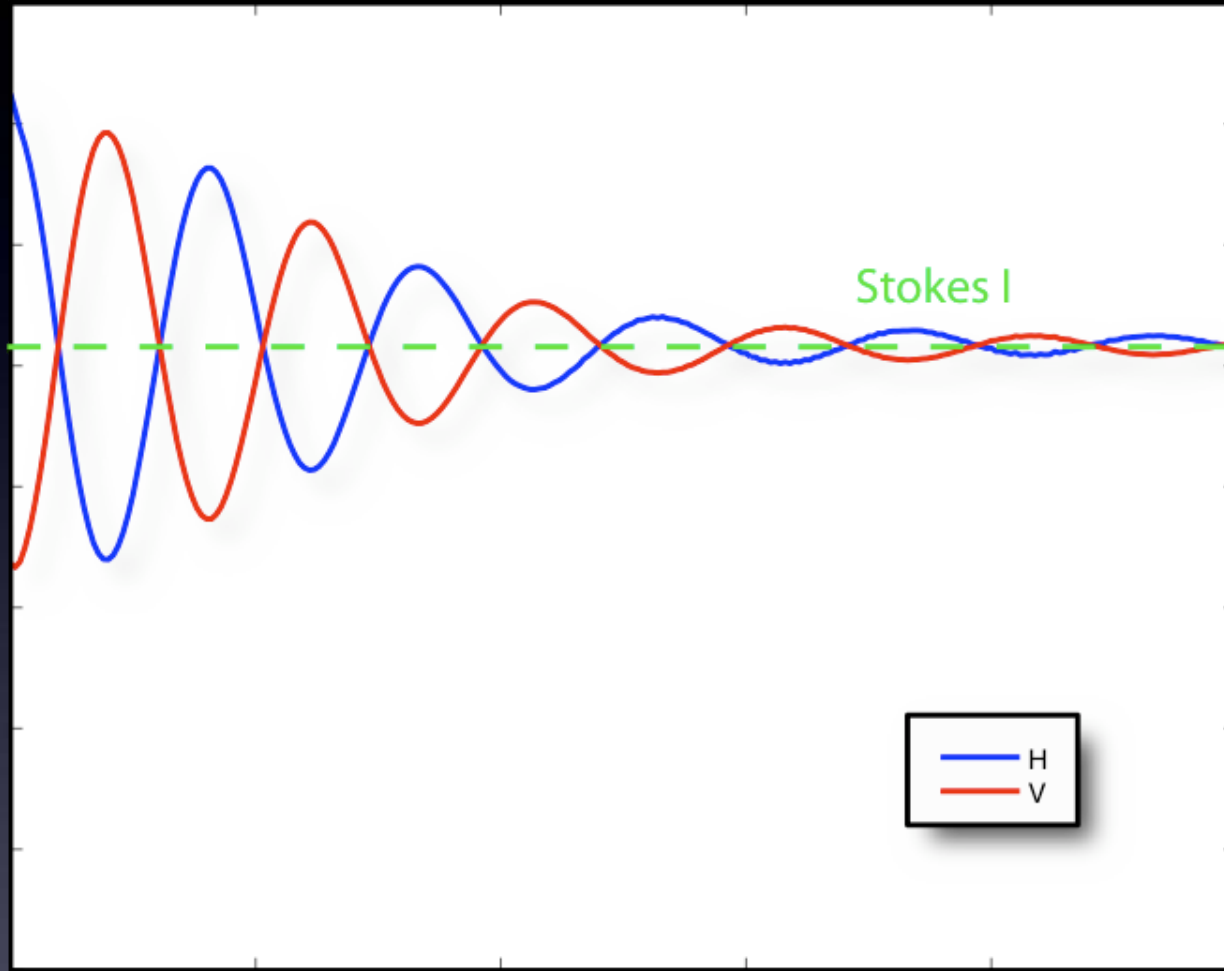


$$Q_{detector} = Q \cos \phi + V \sin \phi$$

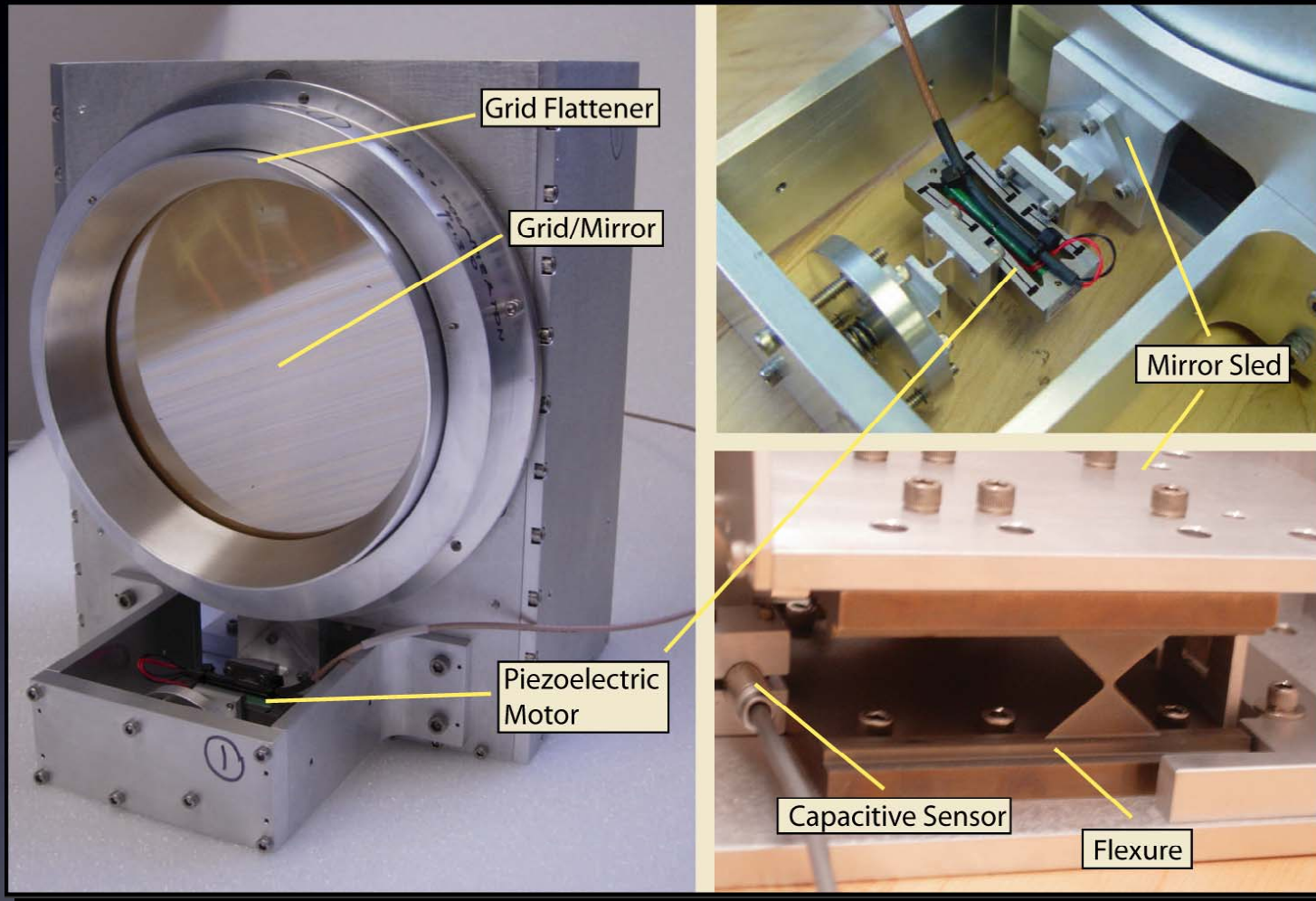




Data Stream



Grid-Mirror Separation

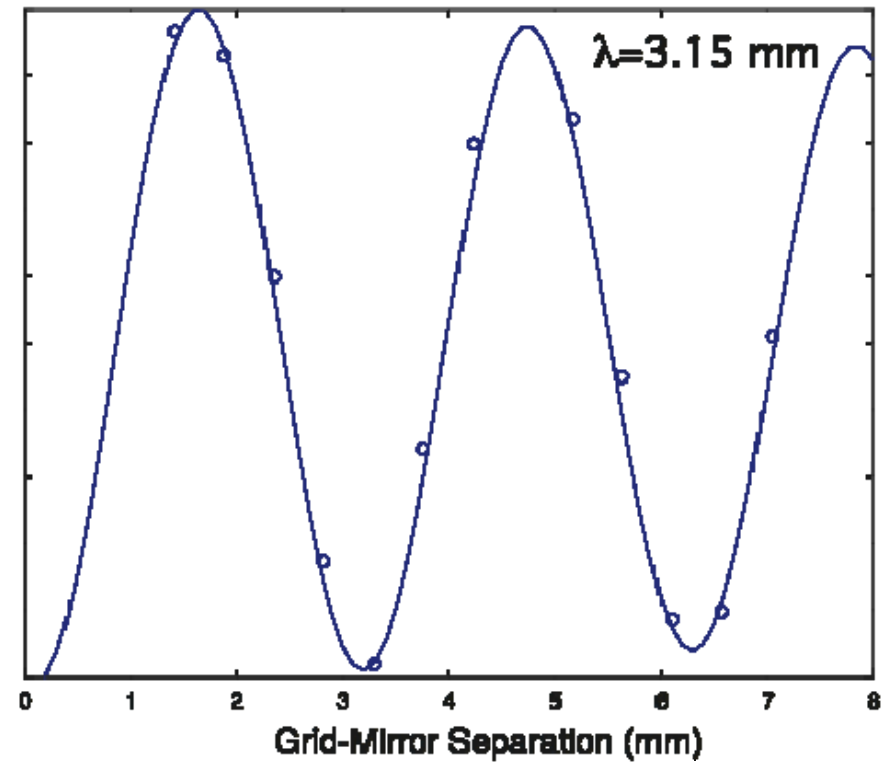
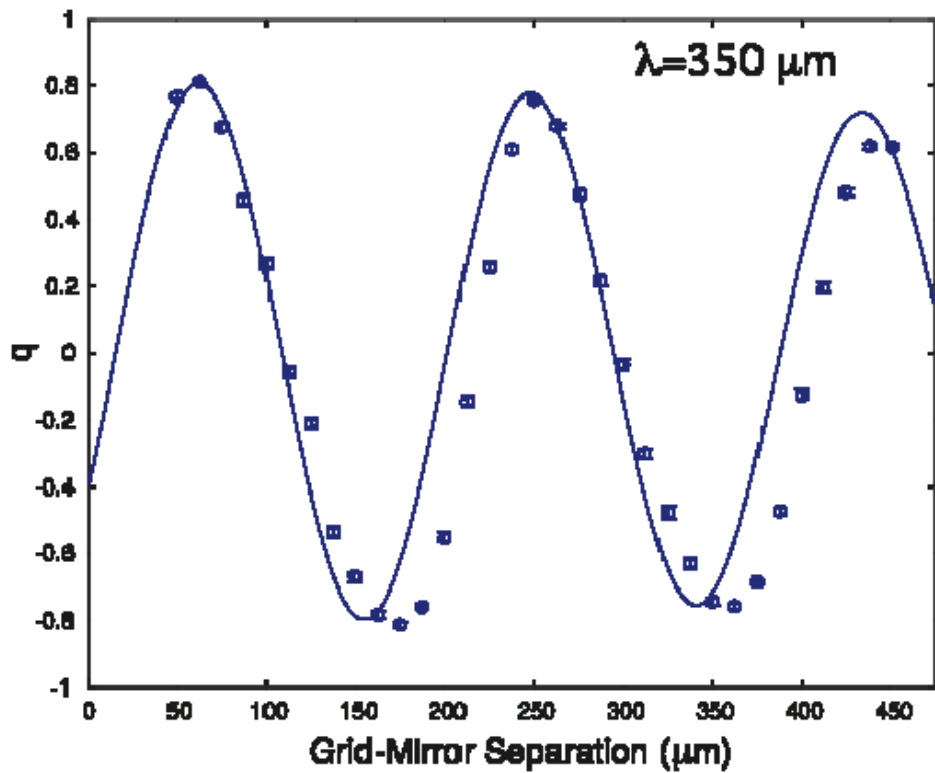


Voellmer et al. (2006), Krejny et al. (2008)



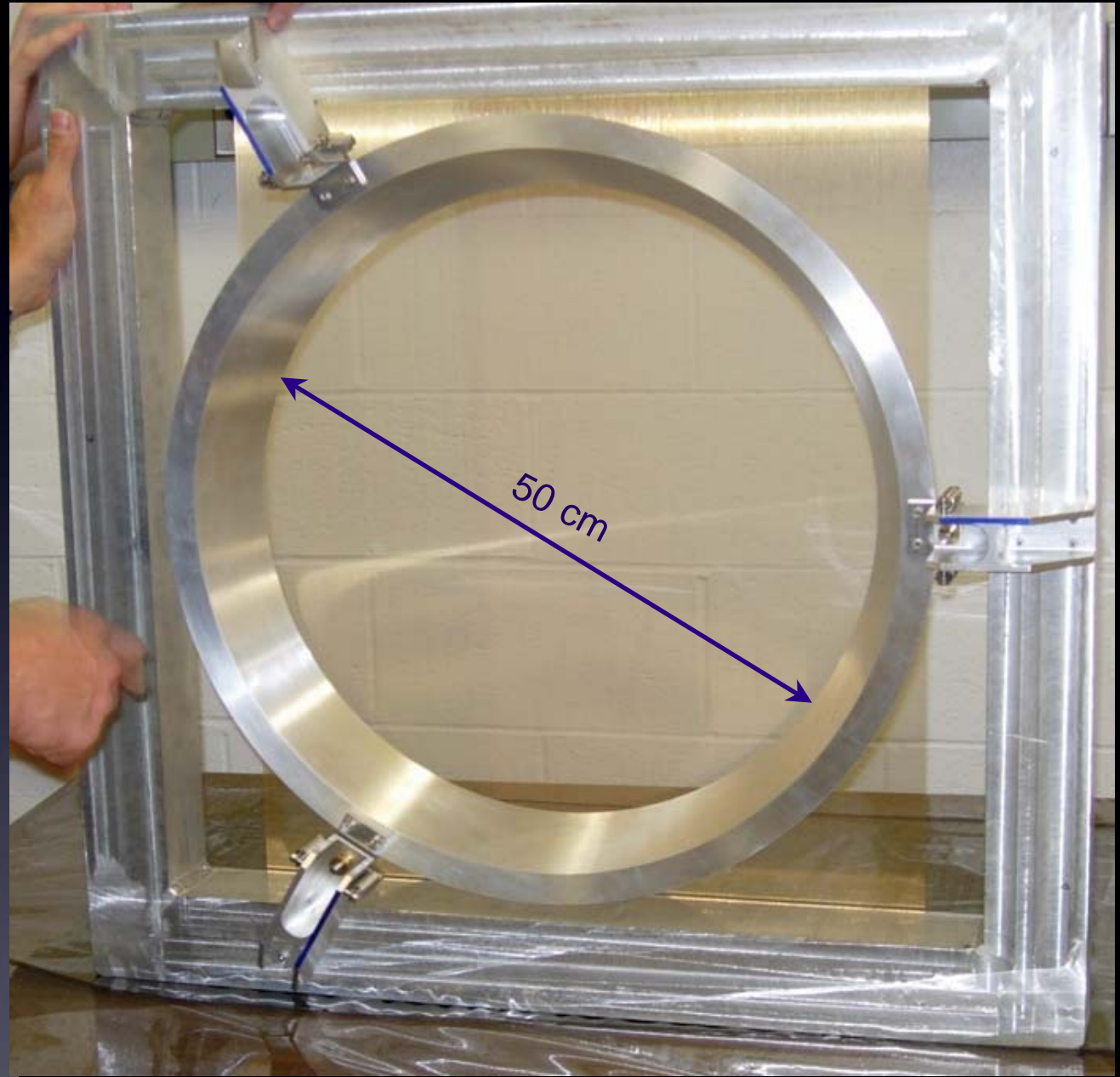
●=350
microns

●=3.15 mm





- Wire diameter = 67 μm
- Wire spacing = 200 μm
- Grid diameter = 50 cm
- Flatness < 50 μm
- Wire resonant frequency > 128 Hz
- ~2 miles of wire
- ~2 Tons of force on the frame





Advantages	Disadvantages
Can be made large for placement close to the sky	Wire is potentially fragile
Interferometric modulation-inherently large bandwidth	Modulation frequency is $\sim < 10$ Hz
Non-dielectric (bi-attenuance, no AR coatings, no deep-dielectric or surface charging)	Lower TRL than competing architectures (3-4)
Good control of Q- \rightarrow U (E- \rightarrow B) mixing (Beam crosspol residuals in V, not U)	Measurement of single Stokes Parameter requires additional time/detectors
Systematics can be placed in modes orthogonal to signal (E.g. variable vignetting/emission is common mode to detectors)	

